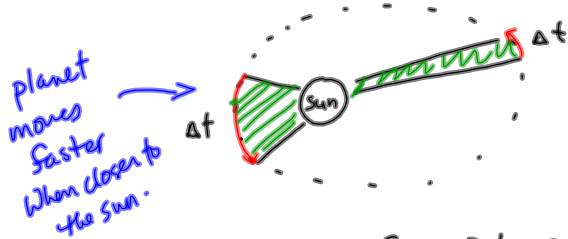


§12-1 Newton's Law of Universal Gravitation

Kepler's Laws

1. Planets move in elliptical orbits.
2. Planets sweep equal areas in equal times:



3. The ratio of r^3/T^2 is constant (for a given central body)

$$K_{sun} = 3.35 \times 10^{18} \frac{m^3}{s^2}$$

$$\frac{r_A^3}{T_A^2} = \frac{r_B^3}{T_B^2}$$

Newton

F_g is directly proportional to m_1 and m_2

F_g is inversely proportional to the square of the separation

$$F_g \propto \frac{m_1 m_2}{r^2}$$

$$F_g = G \frac{m_1 m_2}{r^2}$$

Where m_1 and m_2 are masses (kg)

r is the separation (centre to centre) (m)

F_g is the force of gravity (N)

$$G = 6.67 \times 10^{-11} \frac{N \cdot m^2}{kg^2}$$

See p955

Newton's Hypothesis:

Newton said that $F_g = F_c$

$$F_g = \frac{G m_1 m_2}{r^2}$$

$$F_c = \frac{mv^2}{r} = \frac{m 4\pi^2 r}{T^2} = m 4\pi^2 r f^2$$

orbiting mass



MP/585

Find the mass of sun using the Earth's radius of orbit and its period of orbit:

$$T = 365.25d$$

$$r = 1.49 \times 10^{11}m$$

$$M_{sun} = ??$$

Using Newton's Hypothesis:

$$F_g = F_c$$

$$\frac{G M_{sun} \cancel{m_{earth}}}{r^2} = \frac{\cancel{m_{earth}} 4\pi^2 r}{T^2}$$

$$M_{sun} = \frac{4\pi^2 r^3}{G T^2} \leftarrow \text{Kepler's constant for the sun.}$$

$$365.25d \left(\frac{24h}{1d} \right) \left(\frac{3600s}{1h} \right)$$

$$= 31557600s$$

$$M_{sun} = \frac{4\pi^2 (1.49 \times 10^{11}m)^3}{(6.67 \times 10^{-11} \frac{N \cdot m^2}{kg^2}) (31557600s)^2}$$

$$M_{sun} = 1.97 \times 10^{30} kg$$

TODO:

① PP/586

② Planetary Motion Handout/Animation